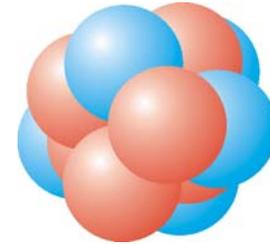


Atomic Nuclei are Made of Protons and Neutrons

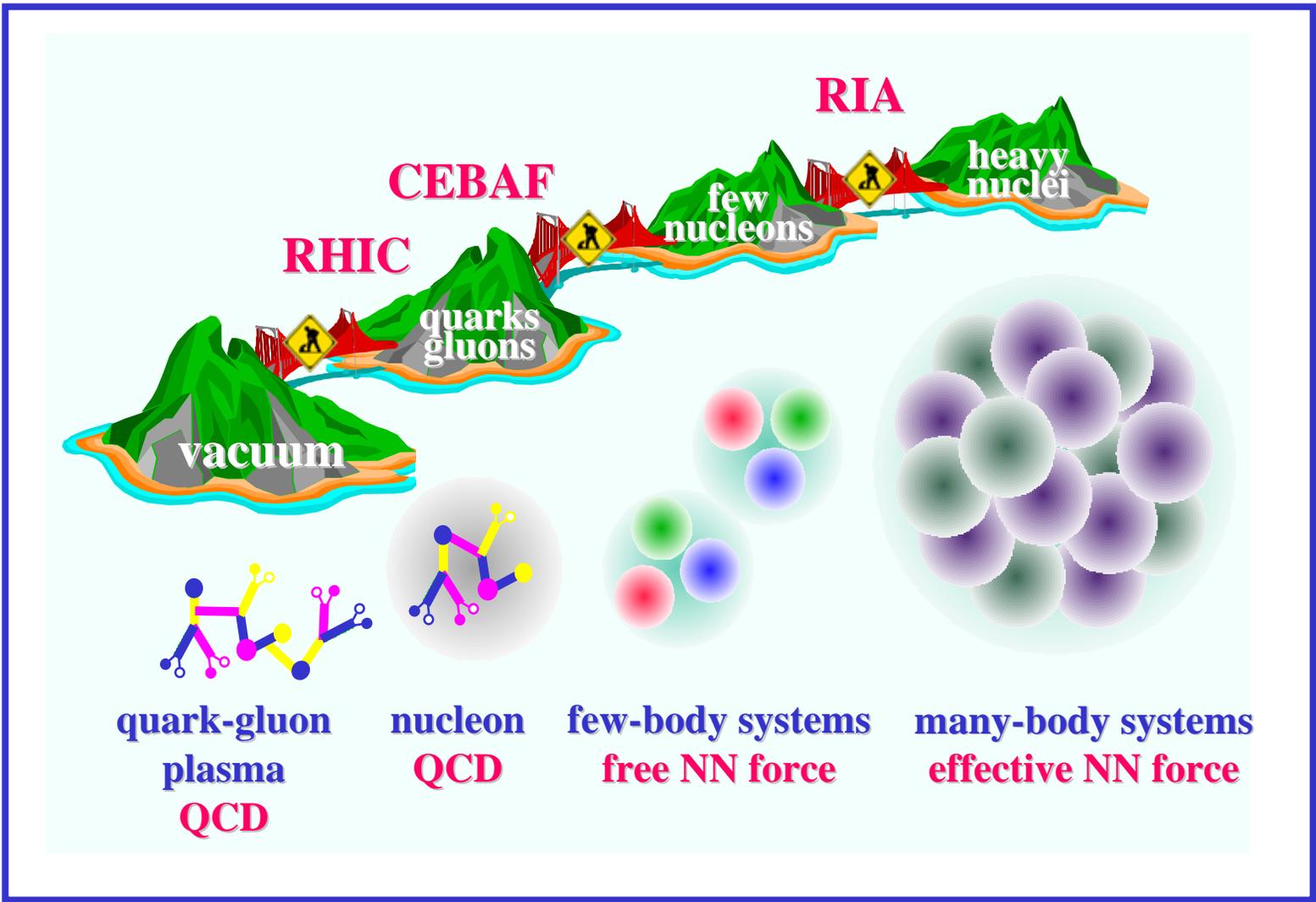
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- The number of protons, Z , determines the chemical element.
- The atomic weight is given by $A = Z + N$, where N is the number of neutrons.
- Fewer than 300 isotopes are stable, all others are unstable (“radioactive”).
- Short-lived isotopes cannot be found on earth -- they have long decayed since earth was formed
- Some radioactive isotopes provide major medical benefits through diagnosis or treatment of diseases; others have important applications, e.g., in biological sciences, environmental sciences, archeology, national security and energy generation.
- We are now able to produce thousands of rare isotopes in the laboratory and explore their properties.



From Simplicity to Complexity

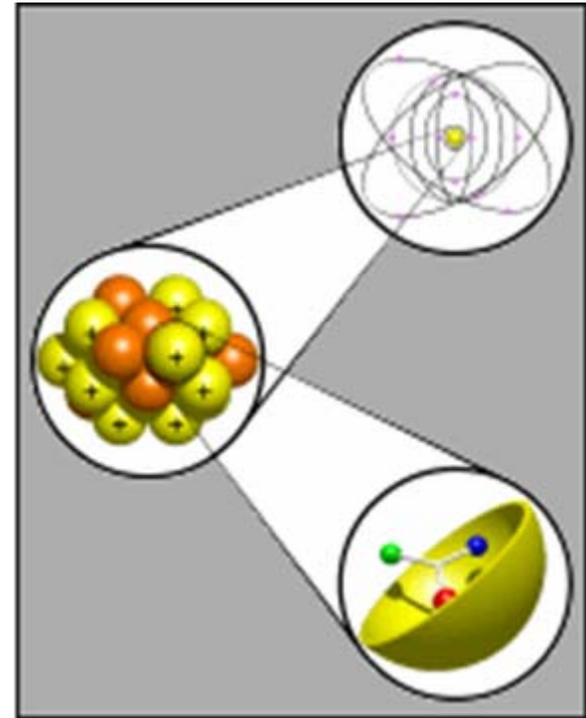
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Thomas Jefferson National Accelerator Facility

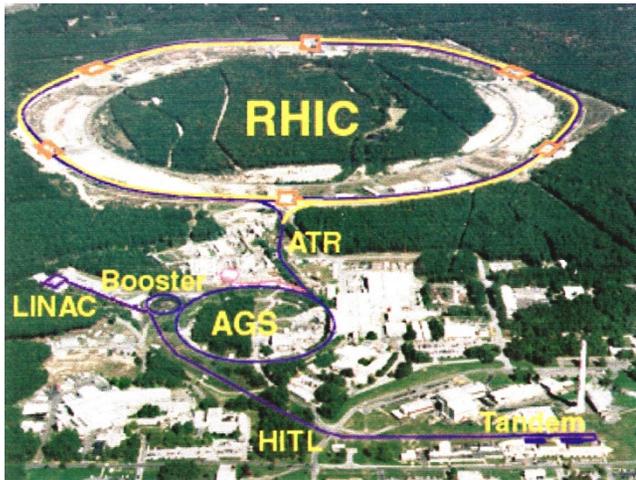
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TJNAF or Jefferson Lab or JLab
or CEBAF = Continuous Electron
Beam Accelerator Facility



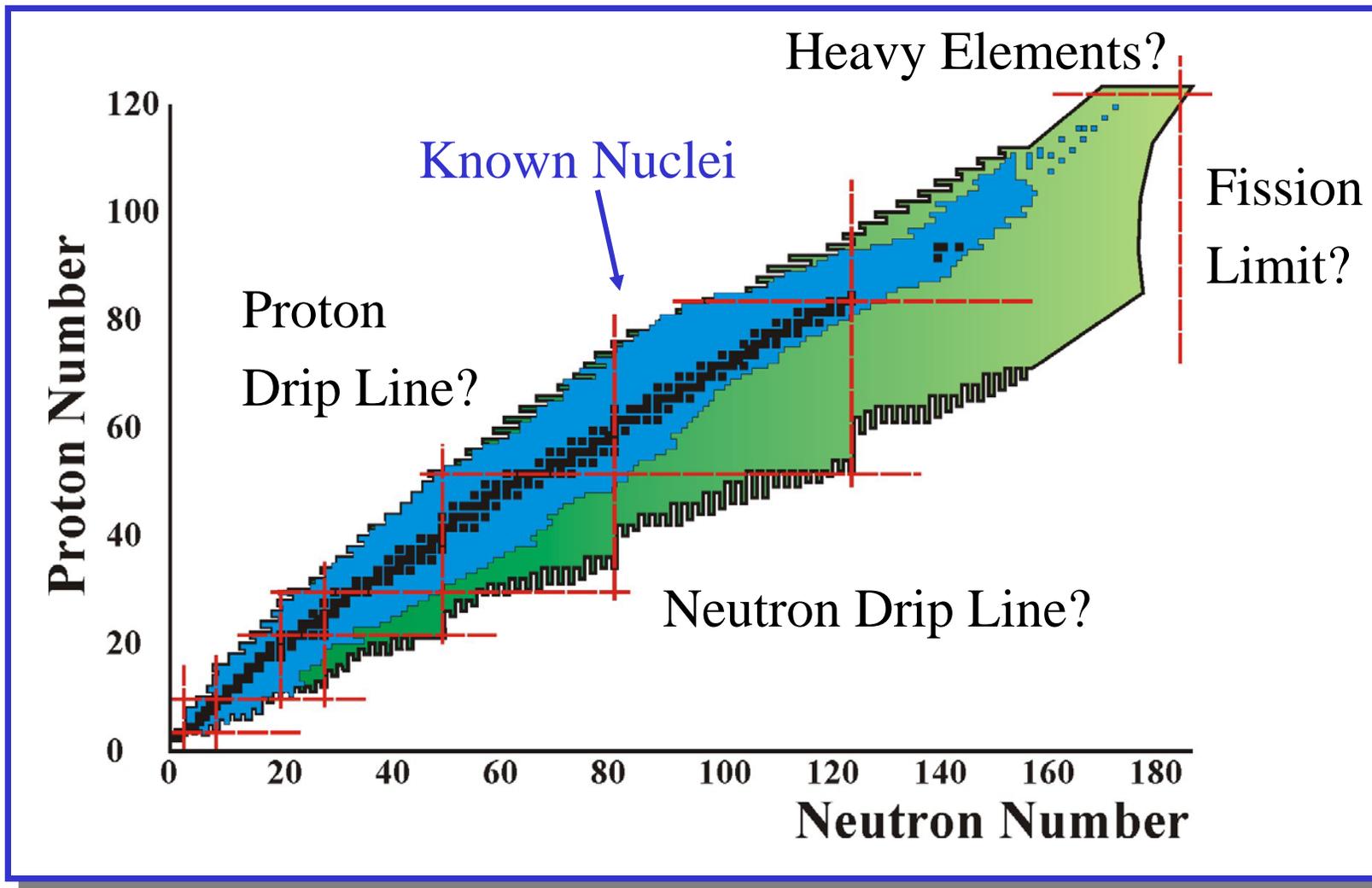
The Relativistic Heavy Ion Collider

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The Chart of the Nuclides

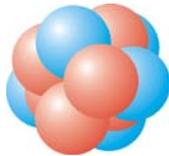
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What is an exotic nucleus?

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Normal Nucleus:



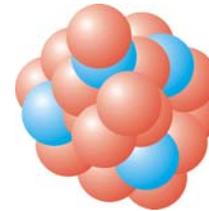
6 neutrons

6 protons (carbon)

^{12}C

Stable, found in nature

Exotic Nucleus:



16 neutrons

6 protons (carbon)

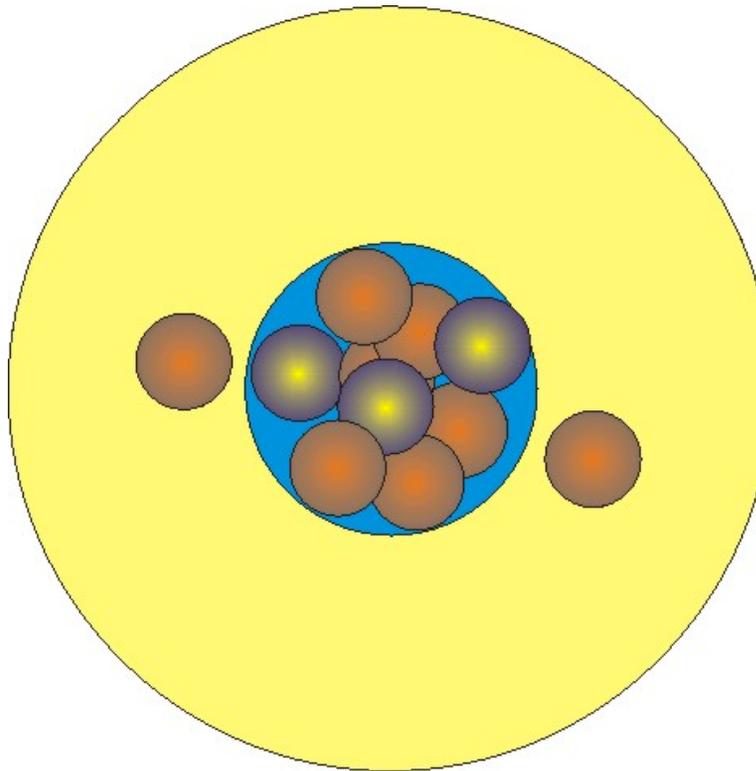
^{22}C

Radioactive, at the limit of nuclear binding

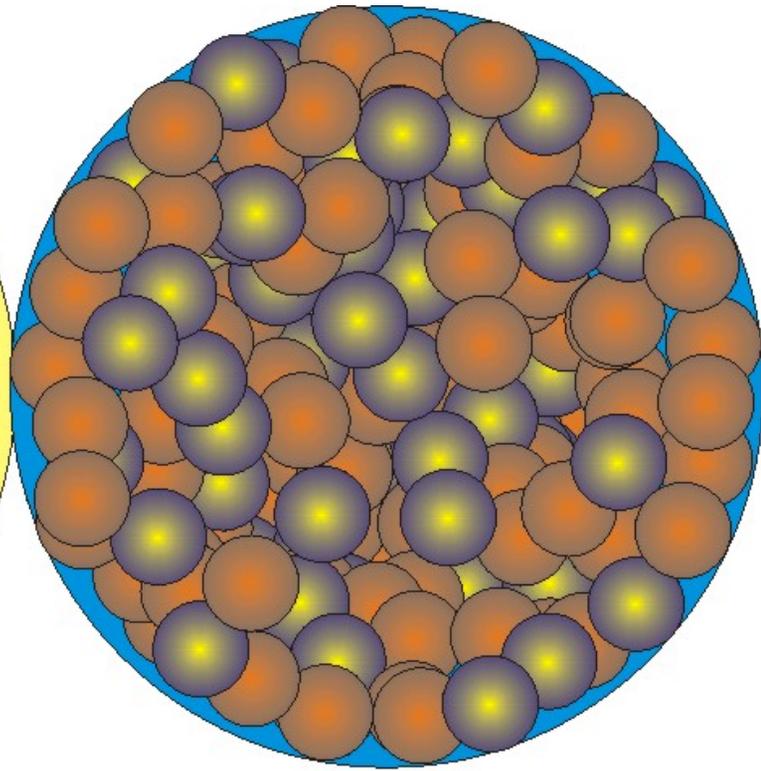
Characteristics of exotic nuclei: Excess of neutrons or protons, short half-life, neutron or proton dominated surface, low binding

Halo Nuclei

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^{11}Li



^{208}Pb



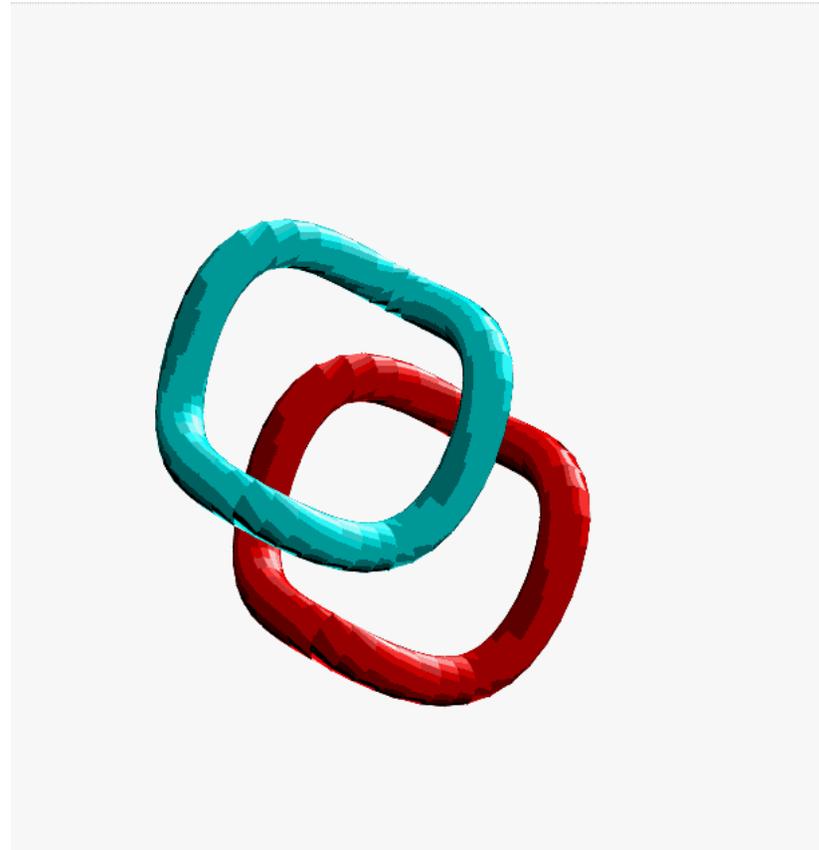
Apollo 17 Crew, NASA

<http://antwrp.gsfc.nasa.gov/apod/ap010204.html>

Borromean Nucleus: ^{11}Li

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Neutron

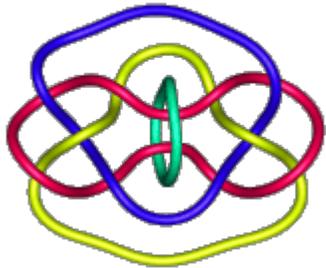


Neutron

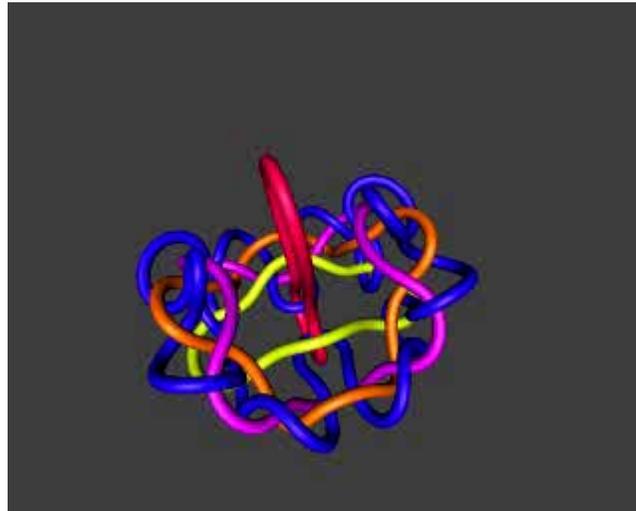
^9Li

Brunnian Links

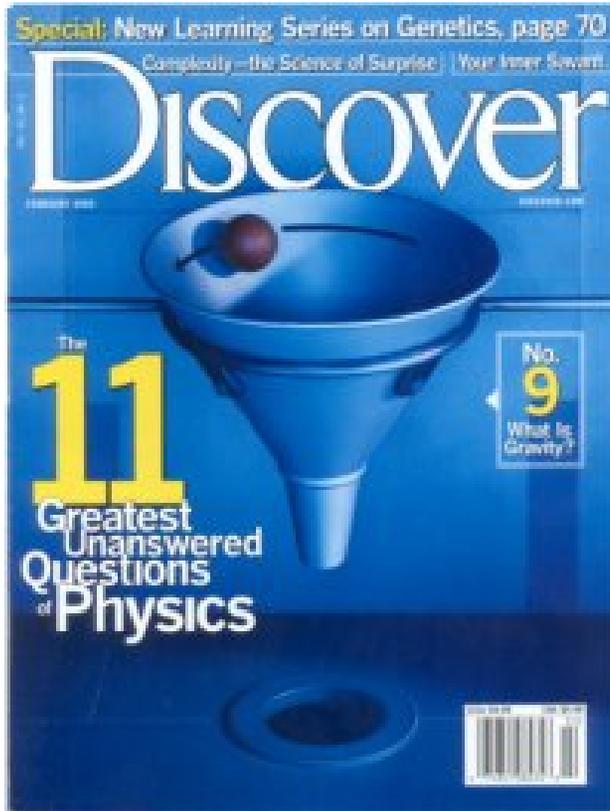
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Robert Scharein
 Department of Computer Science
 University of British Columbia
<http://www.cs.ubc.ca/nest/imager/contributions/scharein/brunnian/brunnian.html>



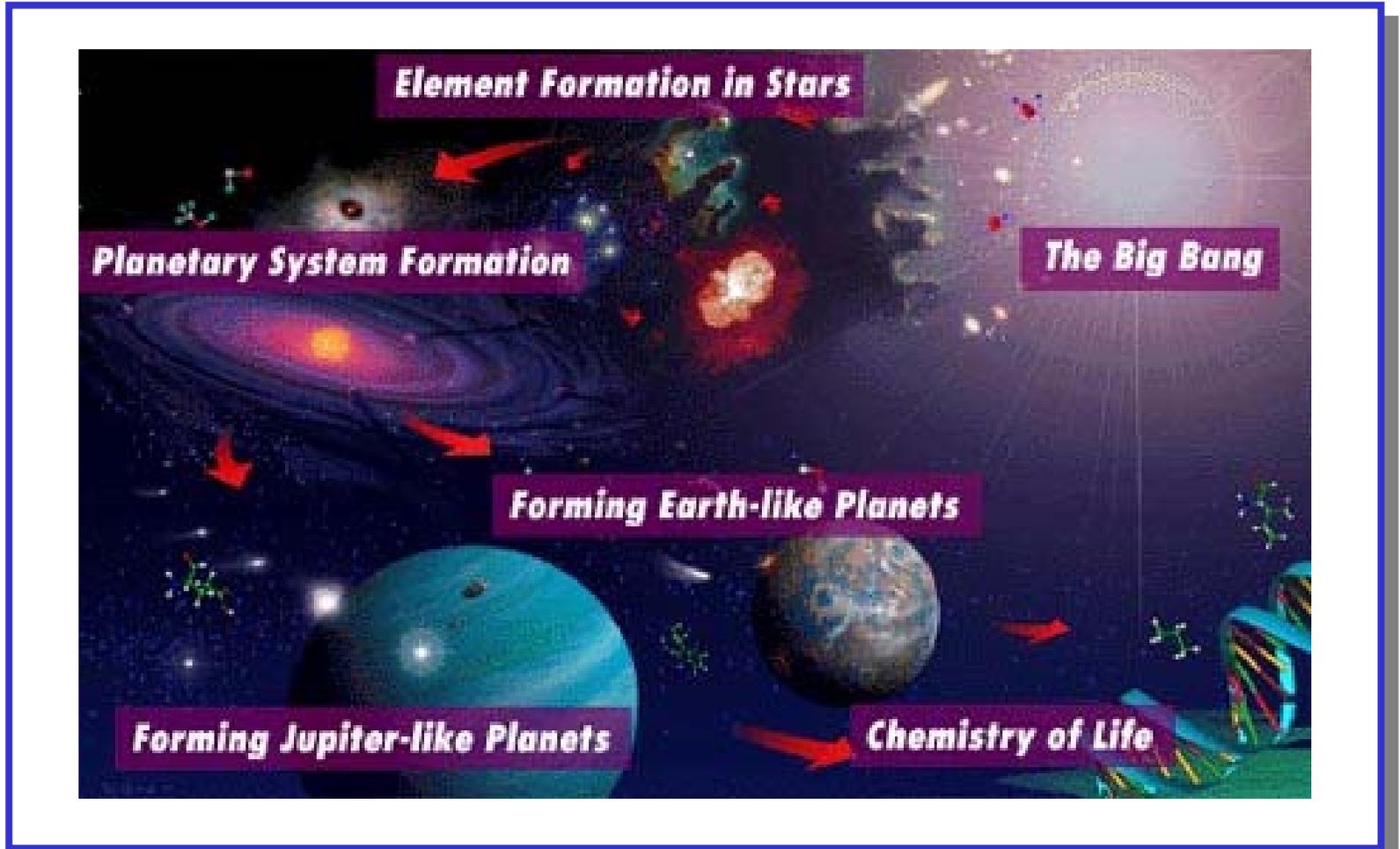
<http://www.cs.ubc.ca/nest/imager/contributions/scharein/brunnian/brun6-rem3.mpg>



1. What is dark matter?
2. What is dark energy?
3. How were the heavy elements from iron to uranium made?
4.

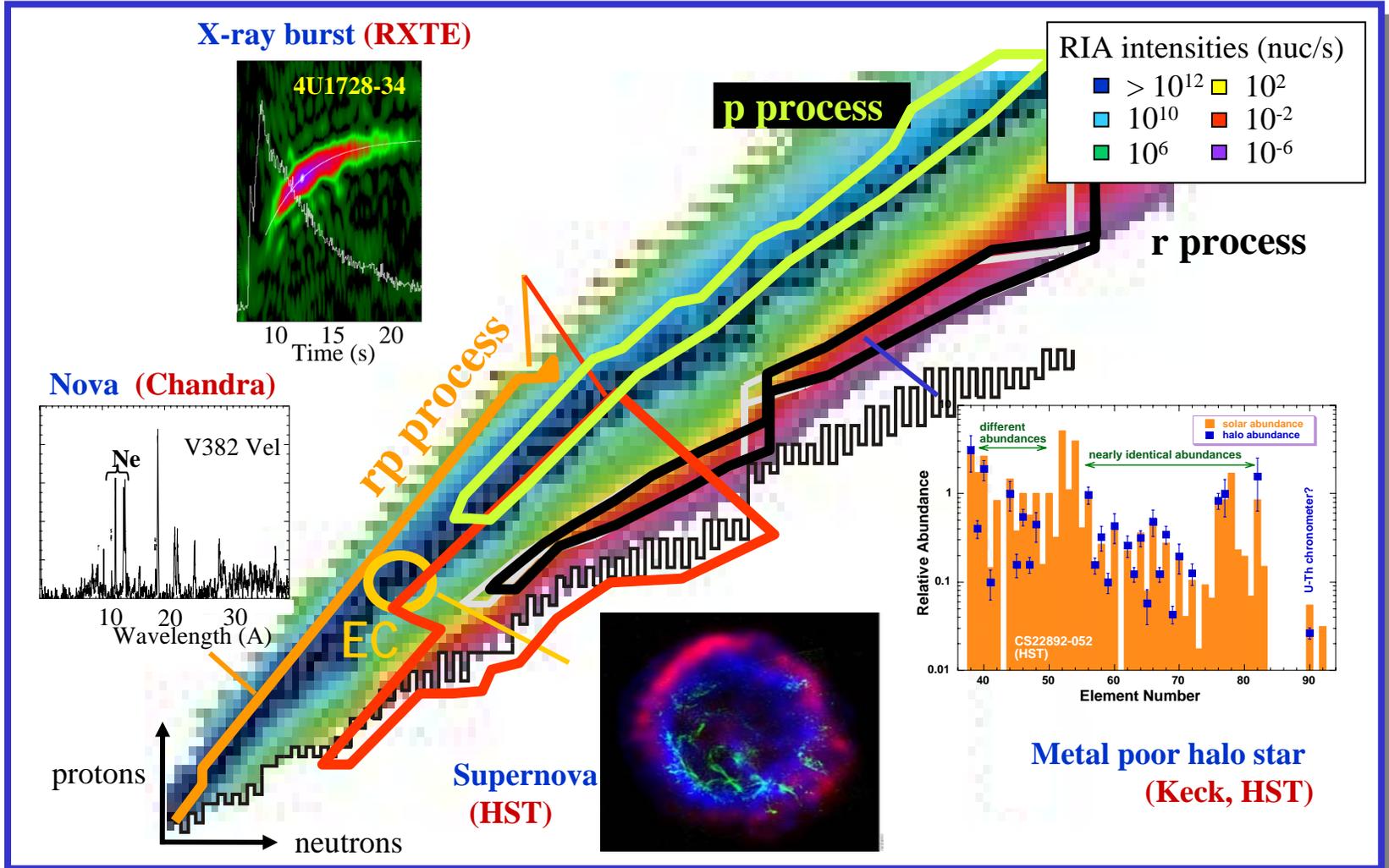
NASA: Timeline of the Universe

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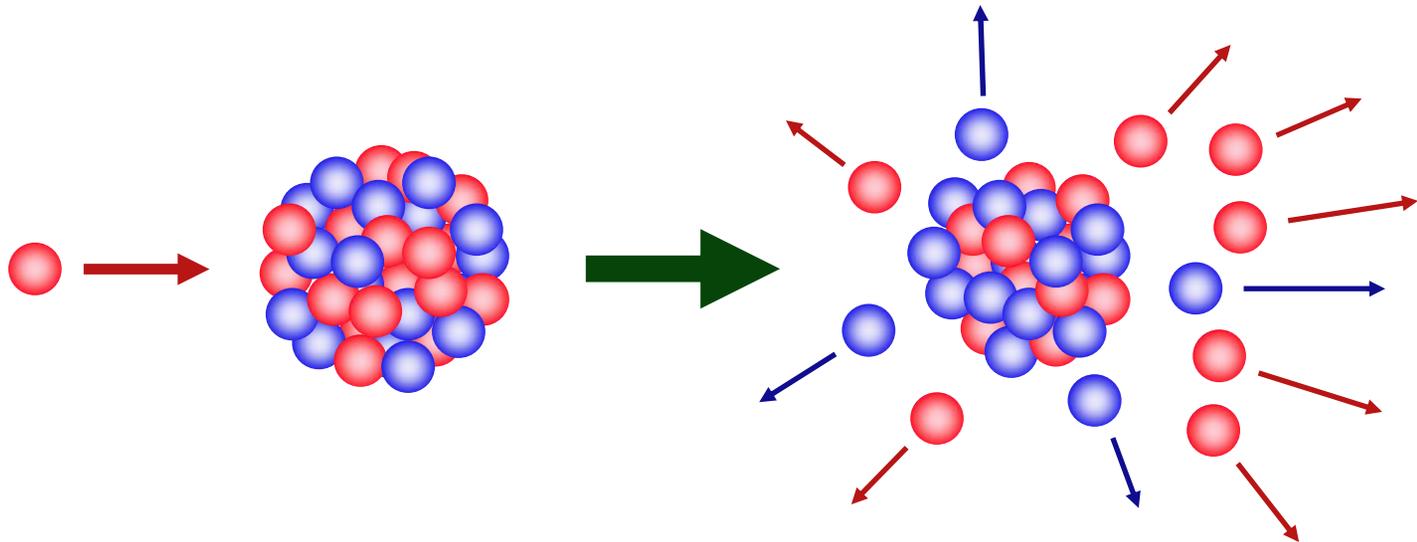


The Origin of the Elements

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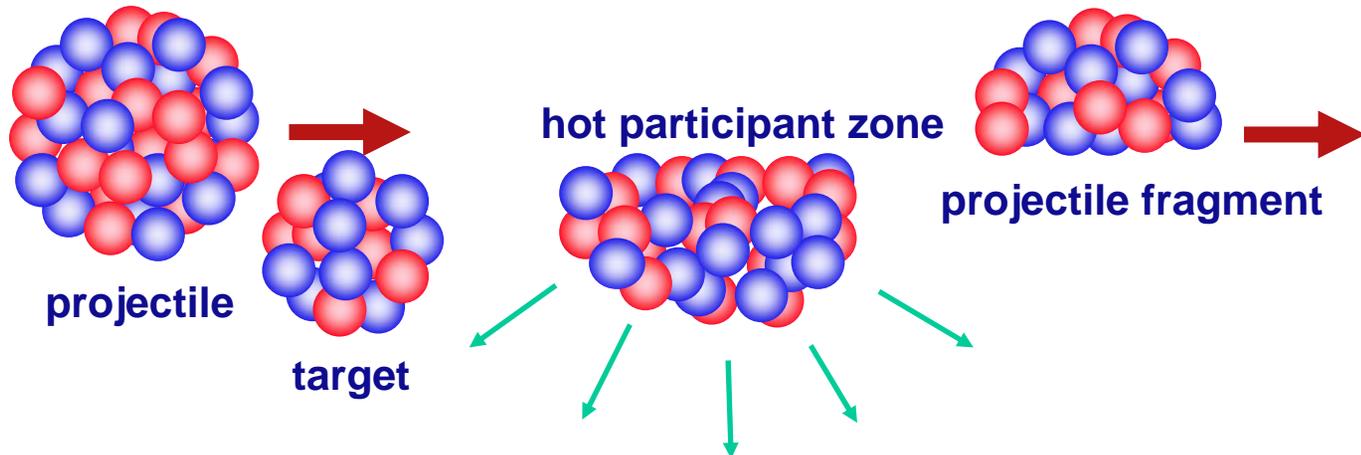


Random removal of protons and neutrons from heavy target nuclei by energetic light projectiles (pre-equilibrium and equilibrium emissions).

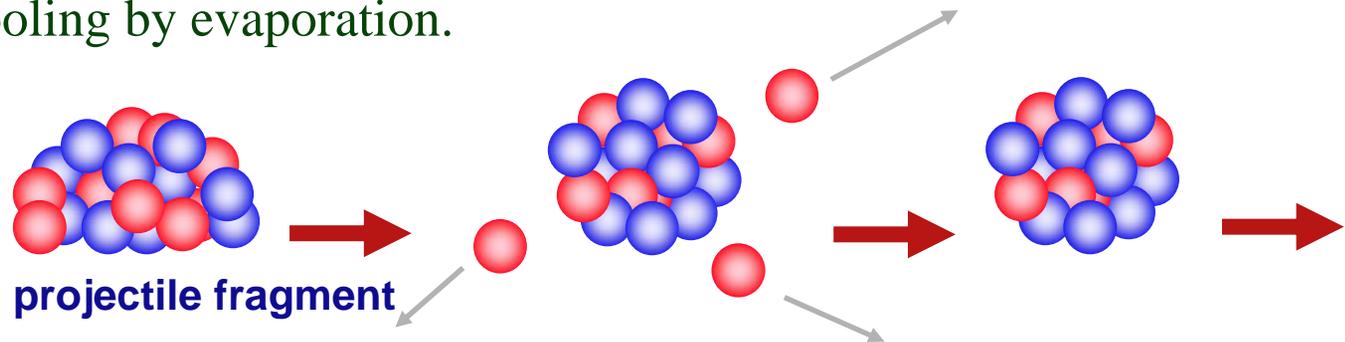


Projectile Fragmentation

Random removal of protons and neutrons from heavy projectile in peripheral collisions



Cooling by evaporation.



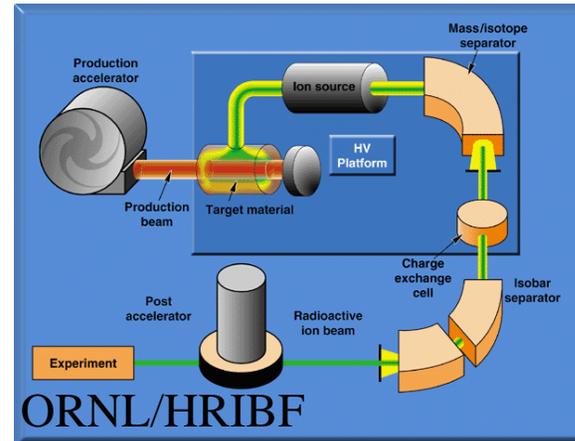
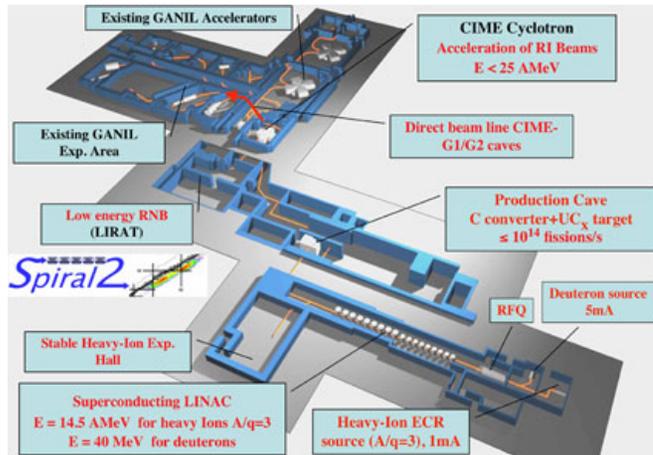
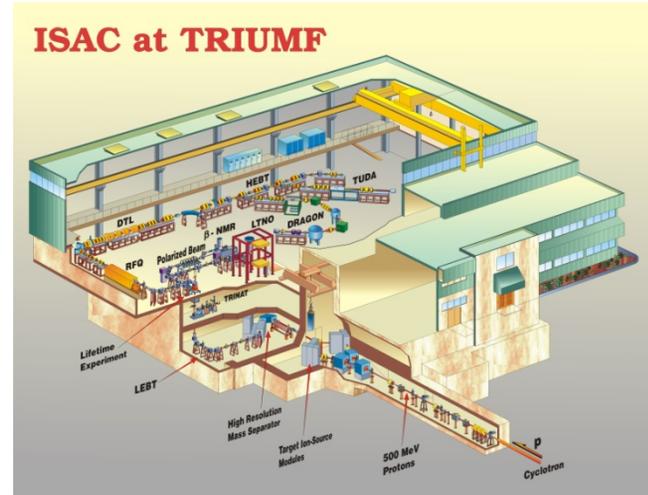
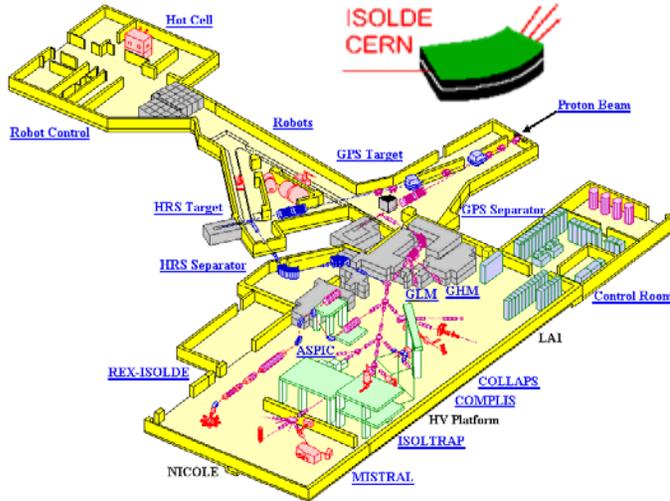
World Wide Effort in RIB Science

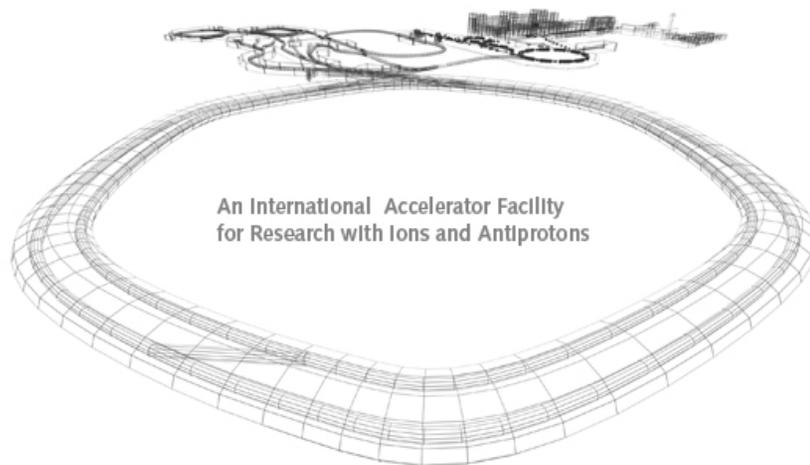
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Plans/Projects at Target Fragmentation Facilities

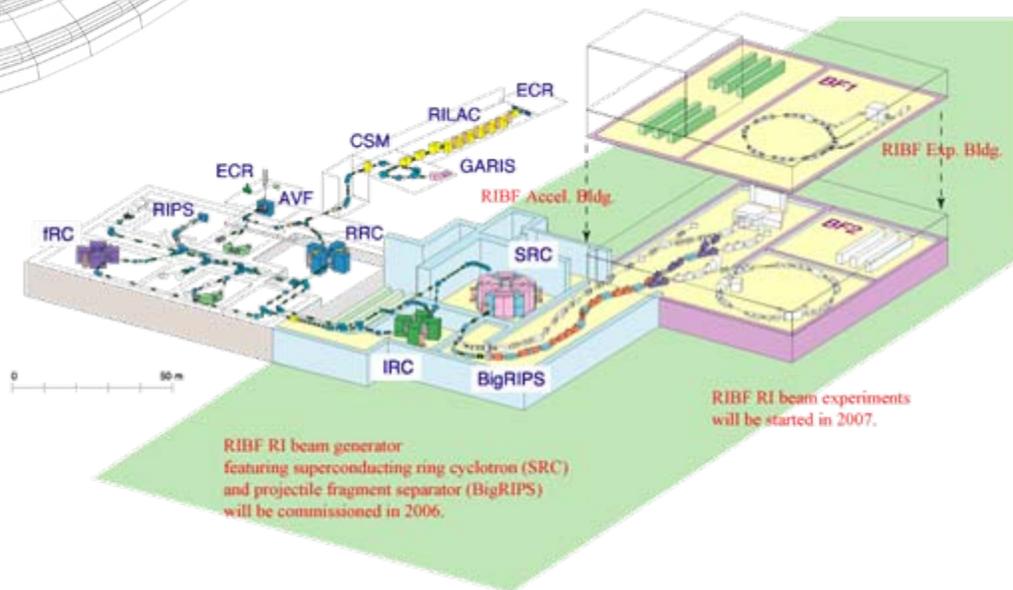
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An International Accelerator Facility
for Research with Ions and Antiprotons

FAIR GSI, Germany



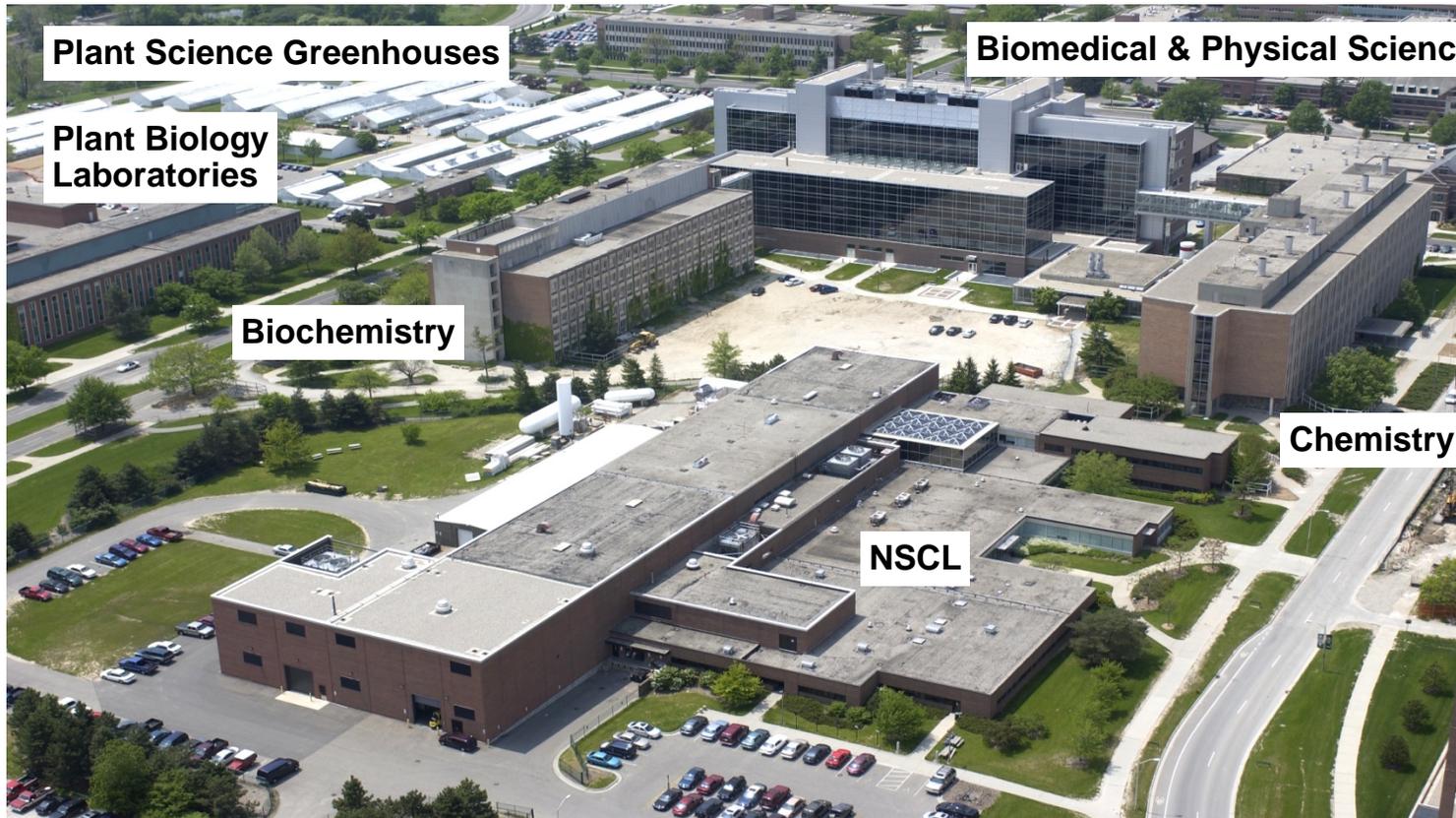
RI Beam Factory RIKEN, Japan



National Superconducting Cyclotron Laboratory



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Nuclear science, astro-nuclear physics, accelerator physics, and societal applications



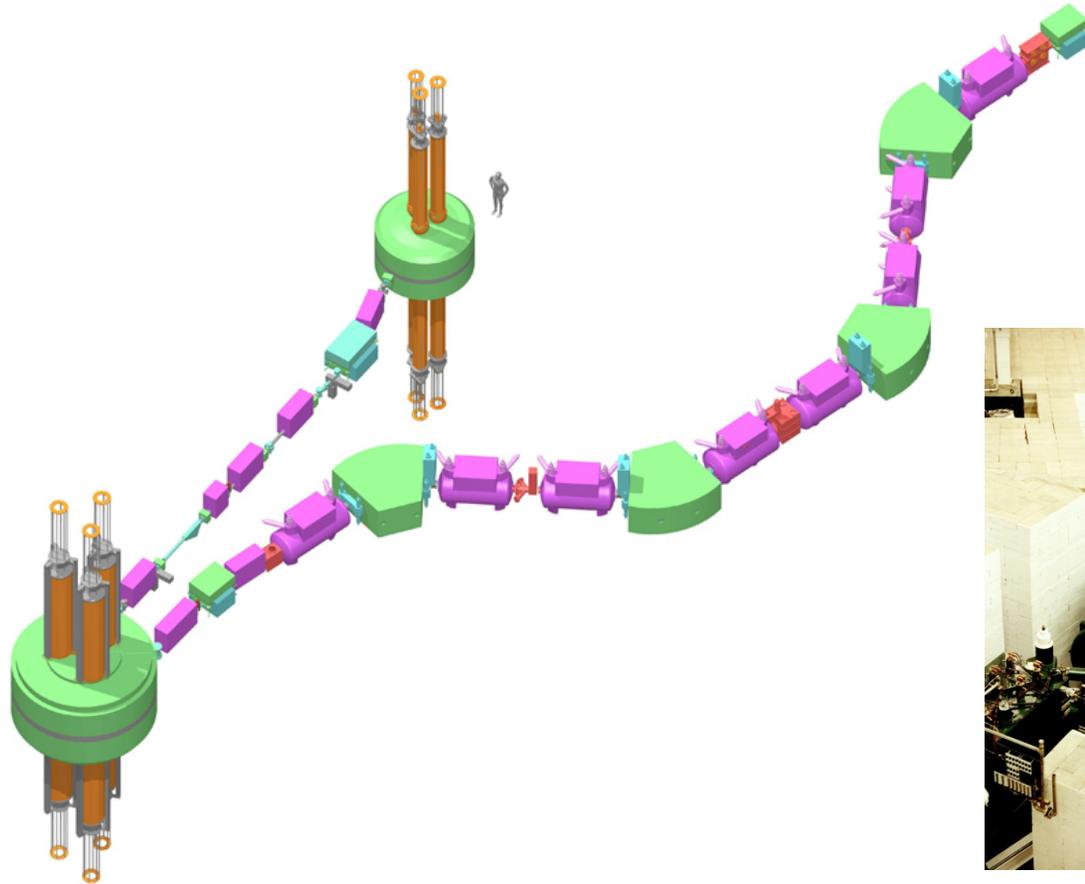
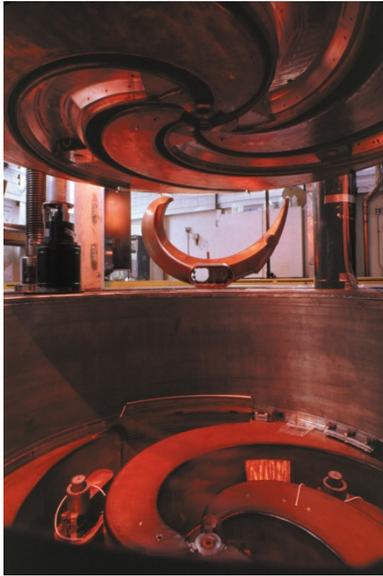
- 295 Employees
 - ▶ Faculty of 29 physicists and chemists
 - ▶ Technical staff of over 150 employees
 - ▶ 52 graduate and 48 undergraduate students
 - ▶ Over 675 separate users to date from over 143 organizations worldwide



- Nuclear physics graduate program ranked #1 in nation where rare isotope research is conducted (#2 overall)

Coupled Cyclotron Facility

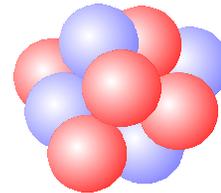
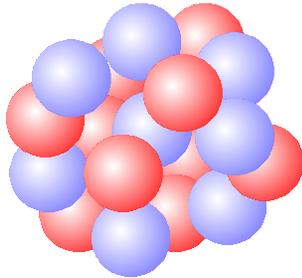
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Fragmentation Reaction

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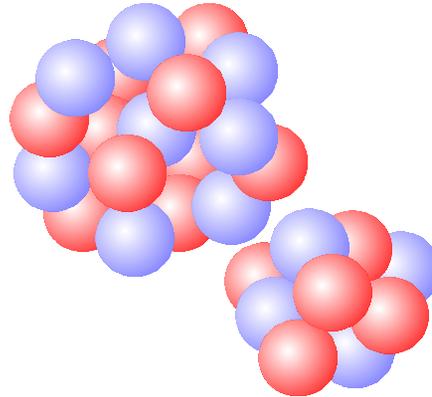
^{18}O beam



^9Be target

$t = -10^{-22}$ sec
 $d = -10$ fm
 .000000000000000039''

^{18}O



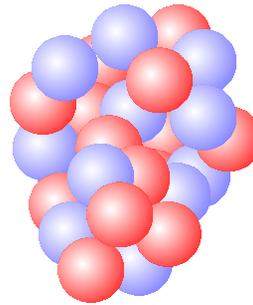
^9Be

$t = -5 \times 10^{-23}$ sec
 $d = -5$ fm

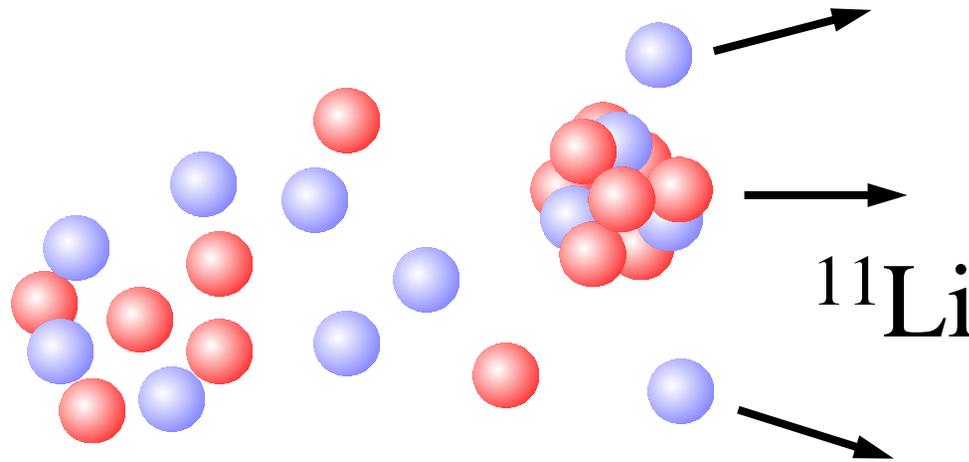
80 MeV/nucleon
 40% speed of light
 278,000,000 mph

Production of ^{11}Li

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$t = 0 \text{ sec}$
 $d = 0 \text{ fm}$



$t = 10^{-22} \text{ sec}$
 $d = 10 \text{ fm}$

^{11}Li

Definitions/Numbers

1pnA, 80 MeV/nucleon, ^{18}O , 8^+

Energy

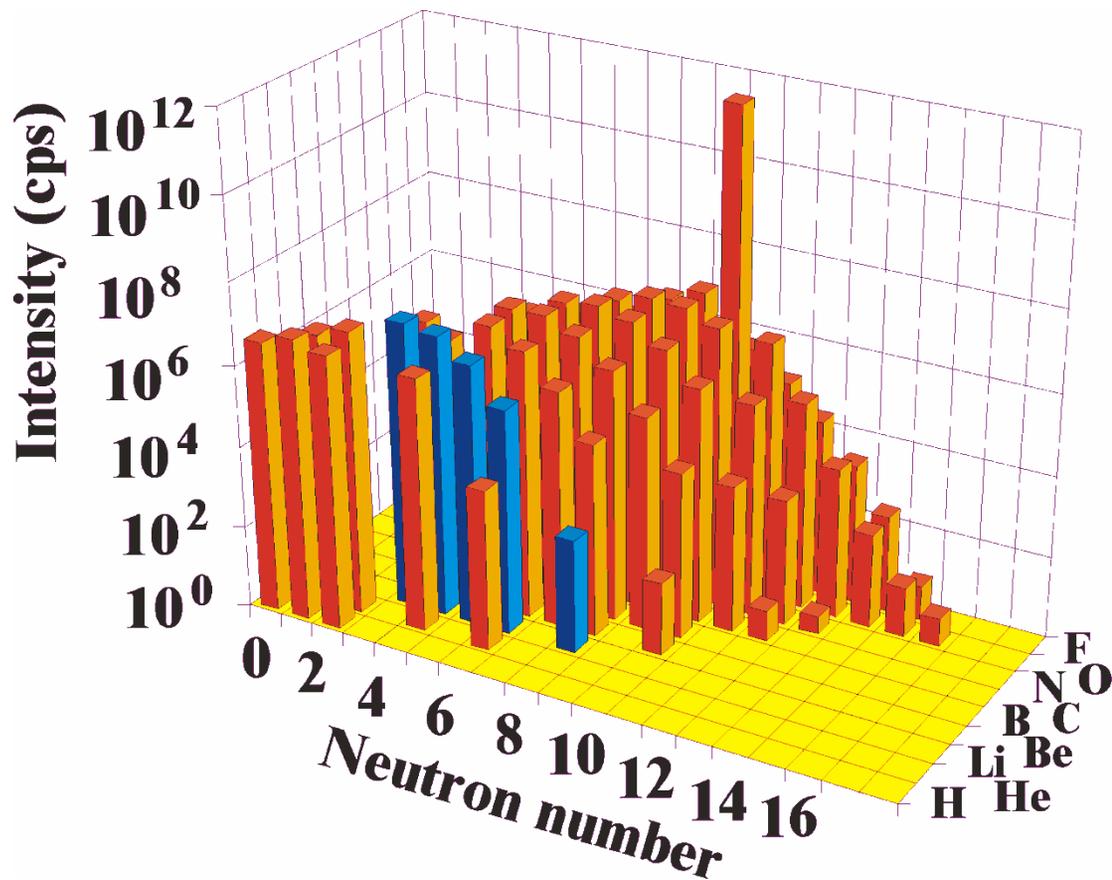
Energy per nucleon:	80 A MeV
Total energy:	1440 MeV
Momentum:	7096 MeV/c
Velocity:	11.7 cm/ns 0.39 c
Rigidity: (p/q)	2.96 Tm

Beam Intensity

Particle Current:	1pnA
Electrical Current:	8enA
Particles:	$6.25 \times 10^9/\text{s}$
Power:	1.44W

Production of Fragments

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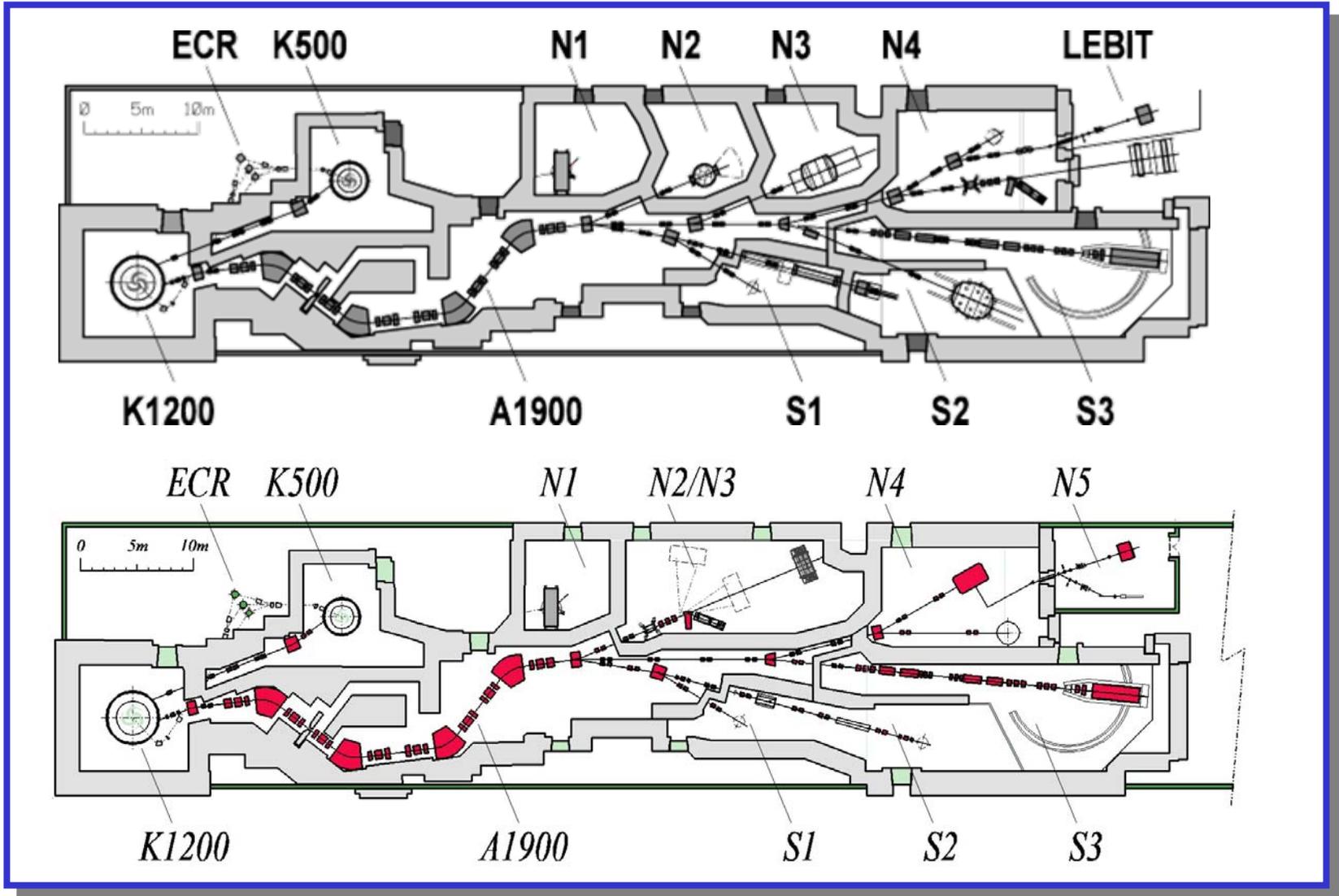


~10 pA ¹⁸O
80 MeV/nucleon

~100 ¹¹Li or
~1/10⁹ ¹¹Li/¹⁸O

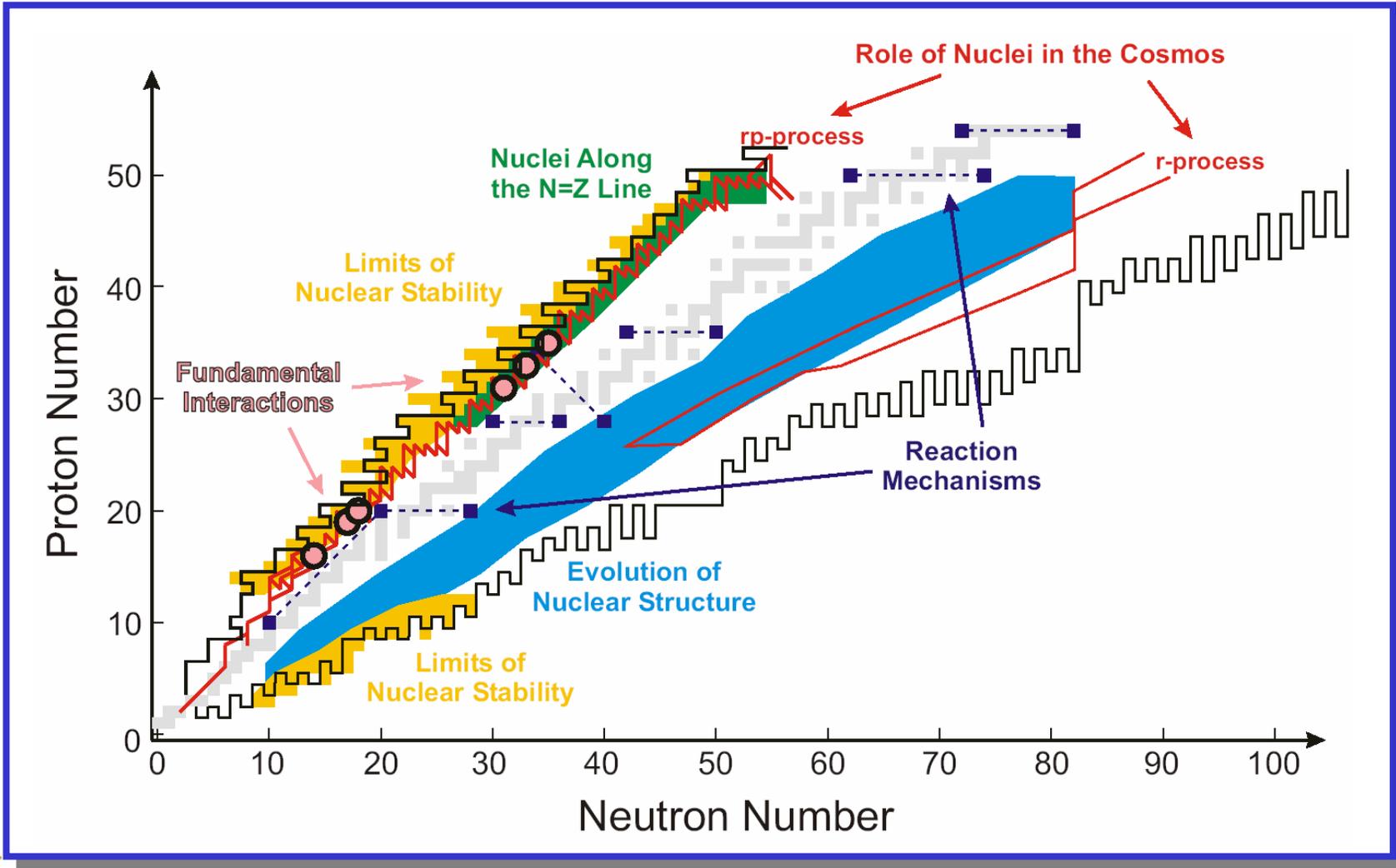
Old and New NSCL Lay-out

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Scientific Reach if the NSCL

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The Future: Department of Energy 20-Year Science Facility Plan



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November 10, 2003 Energy Secretary Spencer Abraham Announces Department of Energy 20-Year Science Facility Plan

Sets Priorities for 28 New, Major Science Research Facilities

WASHINGTON, DC — In a speech at the National Press Club today, U.S. Energy Secretary Spencer Abraham outlined the Department of

RIA
Rare Isotope Accelerator

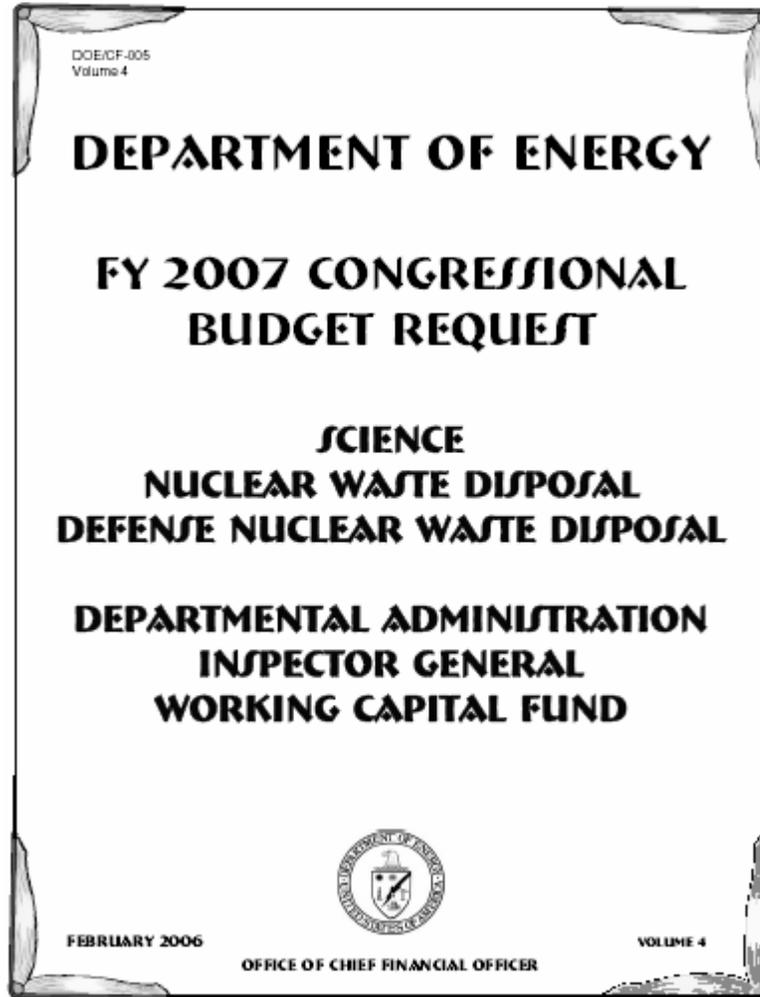




FY 2007 Budget (Feb. 2006)



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Generic R&D in radioactive ion beam development, relevant for **next-generation facilities** in nuclear structure and astrophysics, is supported in FY 2007.

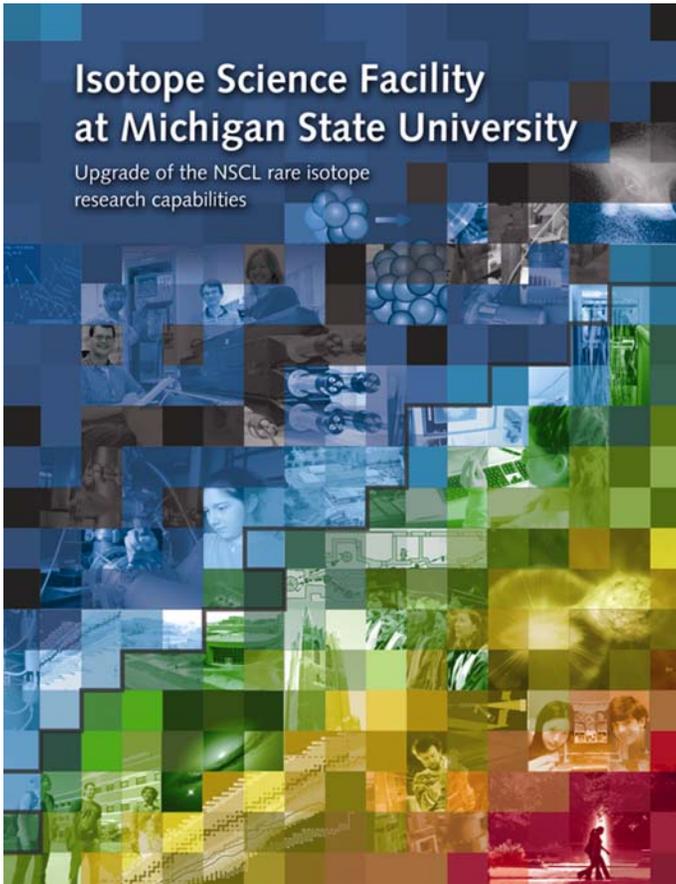


Isotope Science Facility (ISF)

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Isotope Science Facility at Michigan State University

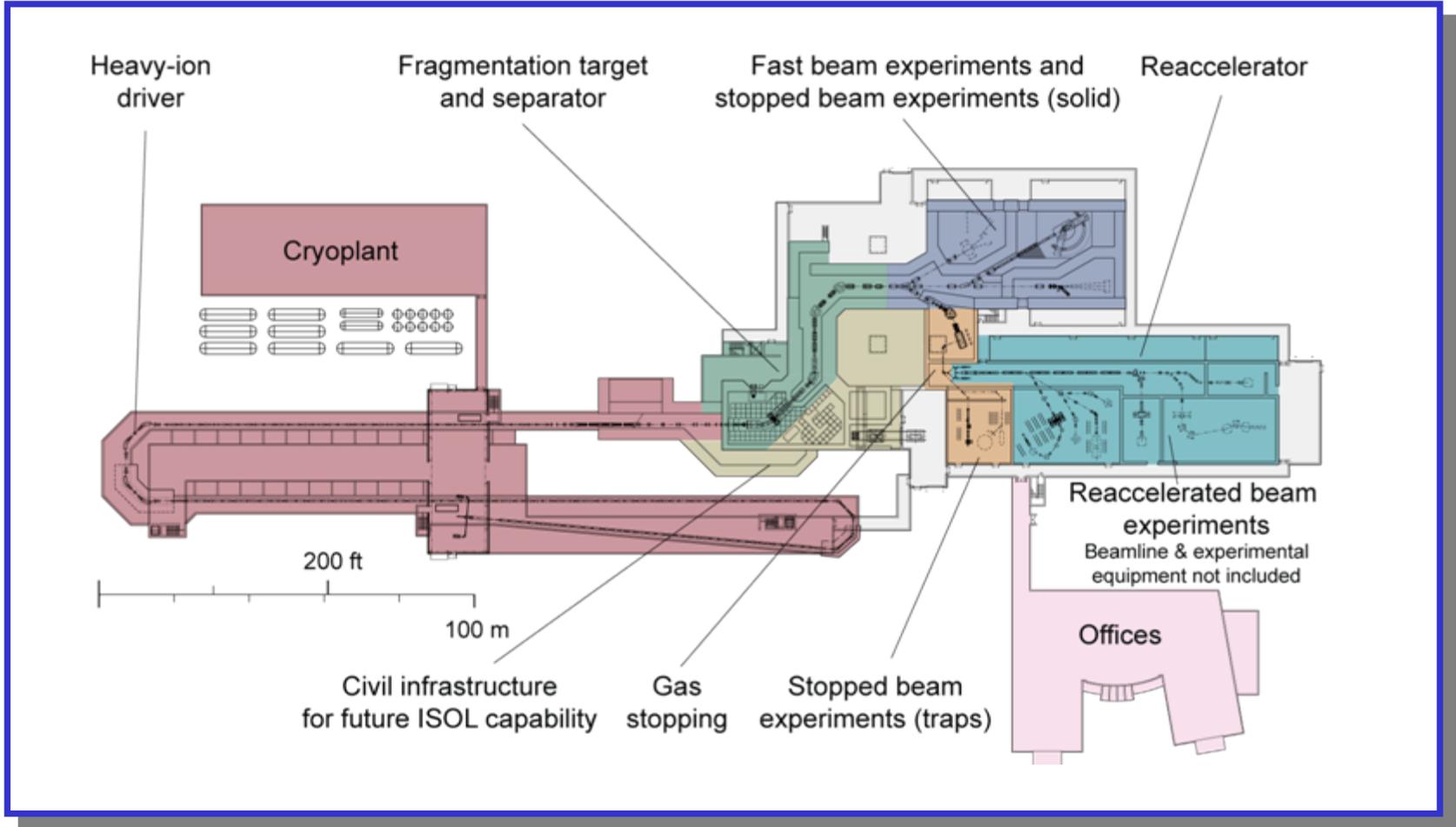
Upgrade of the NSCL rare isotope
research capabilities



- The CCF at the NSCL has a strong rare isotope science program based on in-flight separation in the next decade
- State-of-the-art equipment and efficient operation will keep the NSCL competitive
- The NSCL needs a more powerful driver to ensure world-leading capabilities in the future
- The ISF upgrade is the next logical step that maintains flexible options for science driven upgrades
- The transition from the CCF to the ISF will allow the NSCL to maintain U.S. leadership in the field without disruption

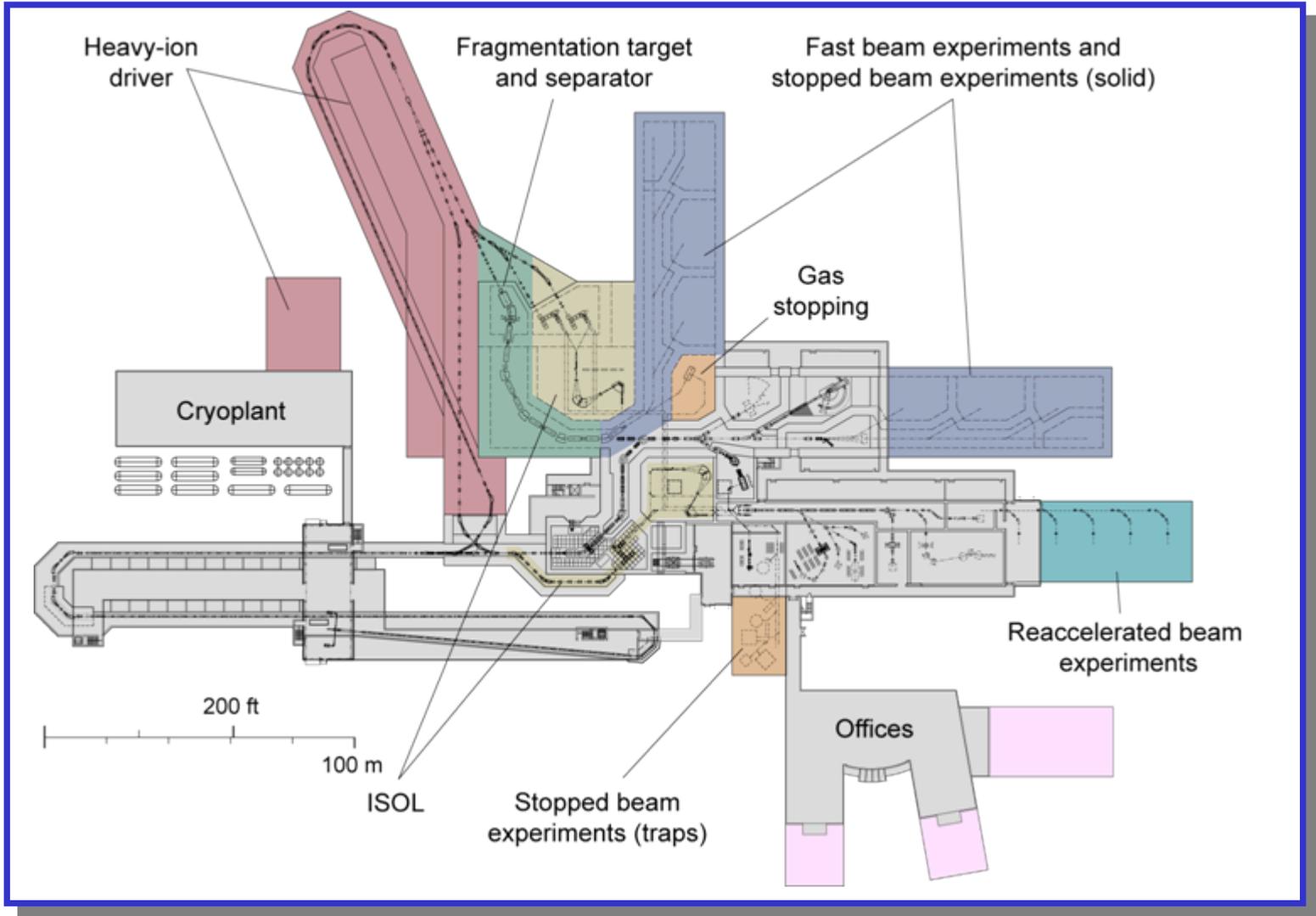
ISF Plan view (South Campus Site)

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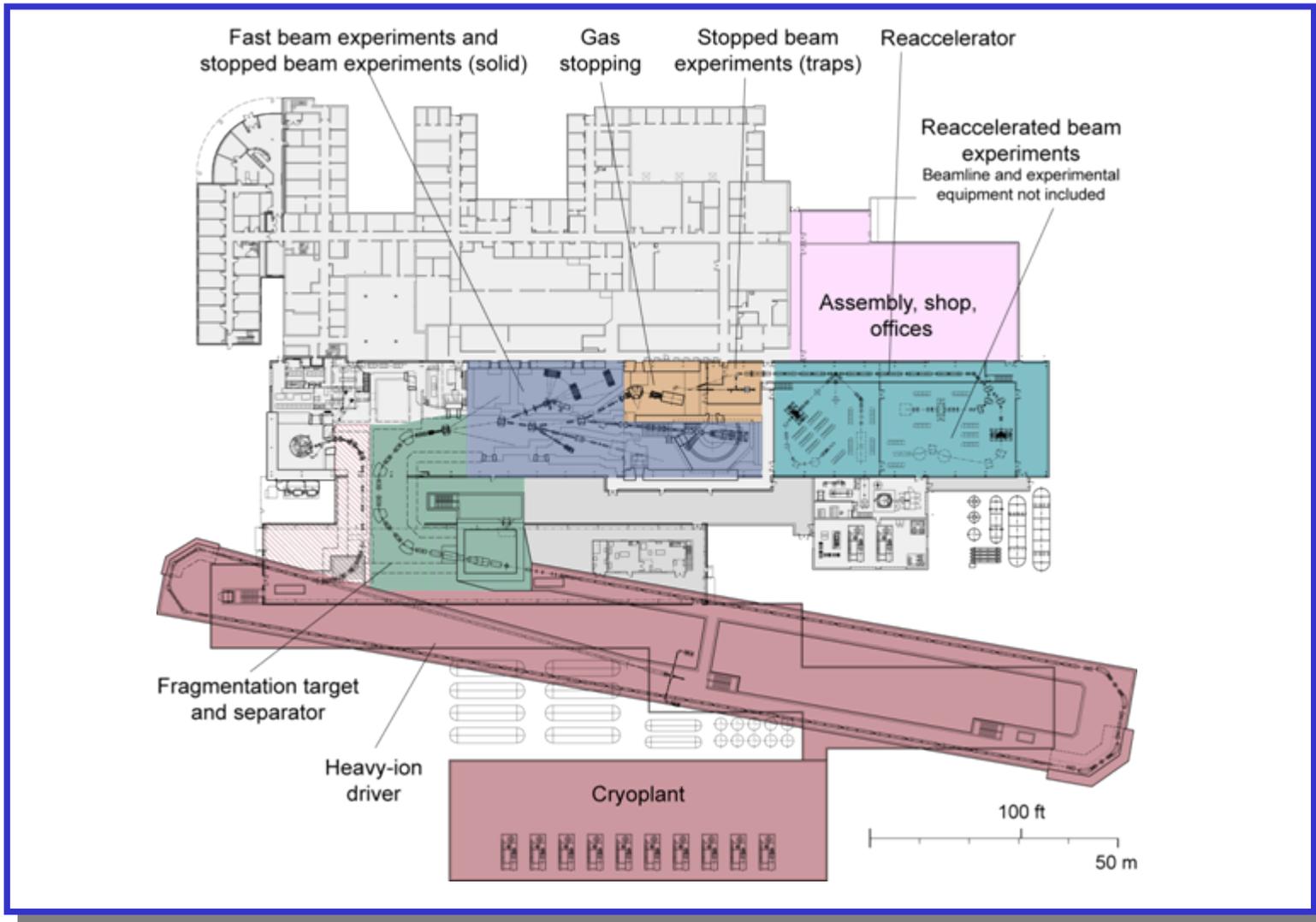
Future ISF Upgrade Options

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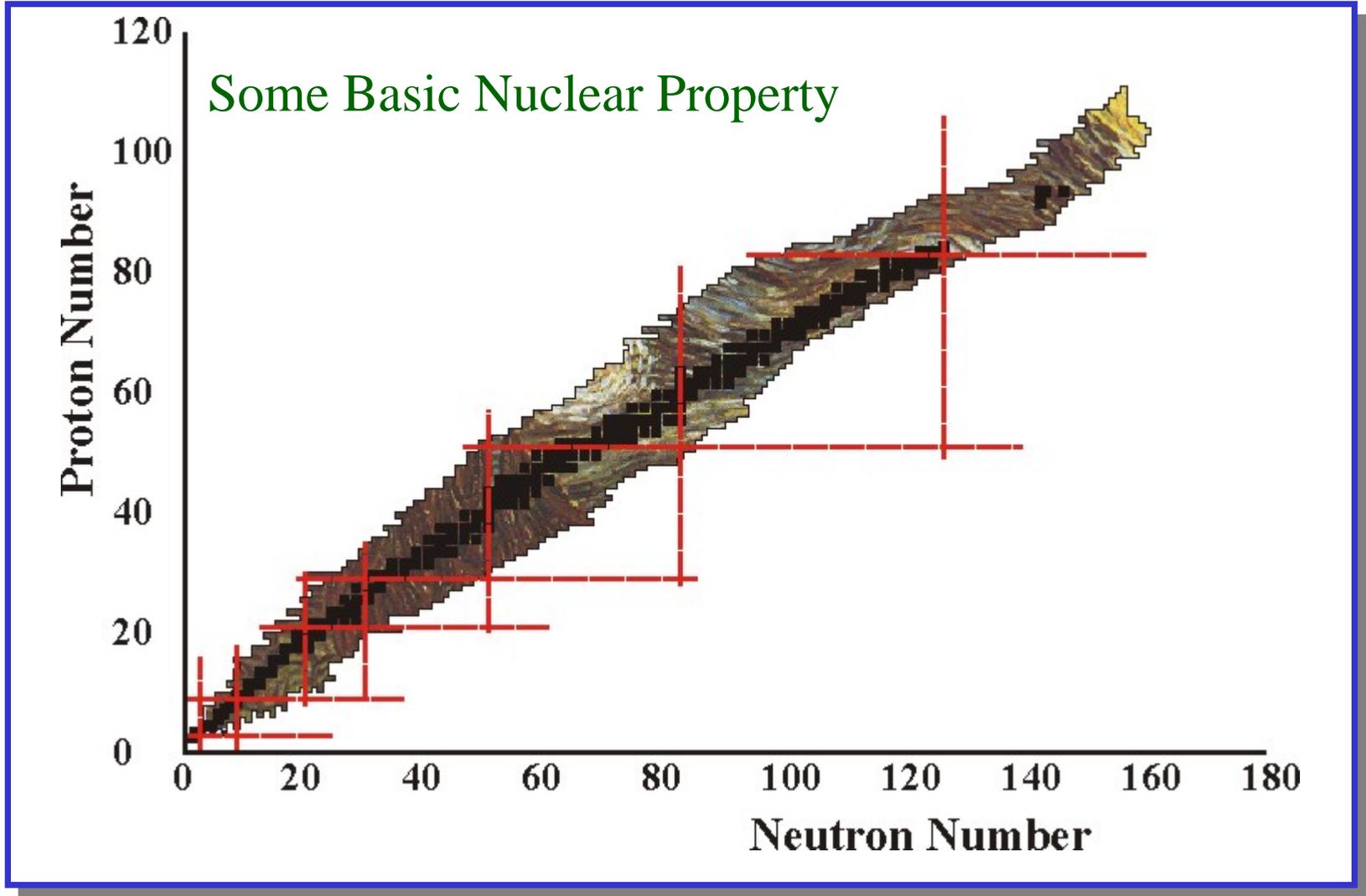
ISF (NSCL Site Option)

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At the moment we are limited in our view of the atomic nucleus

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In the future we will greatly expand our horizons

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